# The Distribution of AGN in Massive Galaxy Clusters at z~l

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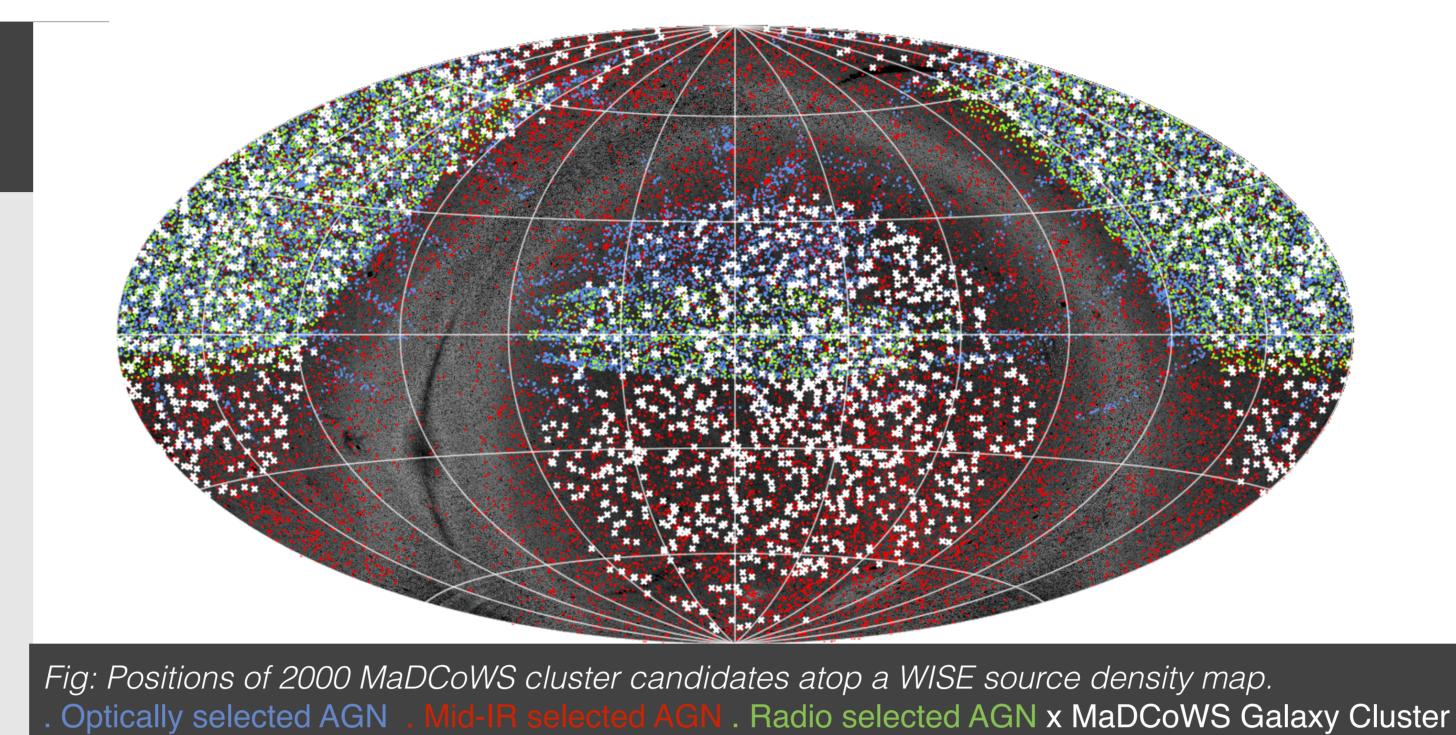
Motivation

Investigate how galaxy cluster environments affect AGN triggering in cluster galaxies and how the feedback of cluster galaxies affect the cluster environment.

## Galaxy Cluster Sample

Massive and Distant Clusters of WISE Survey (MaDCoWS): 2000 galaxy clusters discovered in *WISE* with *Spitzer* followup forthcoming for ~2000.

Largest galaxy cluster sample at z=0.8-1.3.



## AGN Catalogs

Optically-selected AGN:
Richards+2015, 1.9x10<sup>6</sup> AGN

Mid-IR-selected AGN:
Secrest+2015, 1.4x10<sup>6</sup> AGN

Radio AGN: Faint Images of the Radio Sky at Twenty-cm (FIRST), 6.5x10<sup>5</sup> AGN

Methods

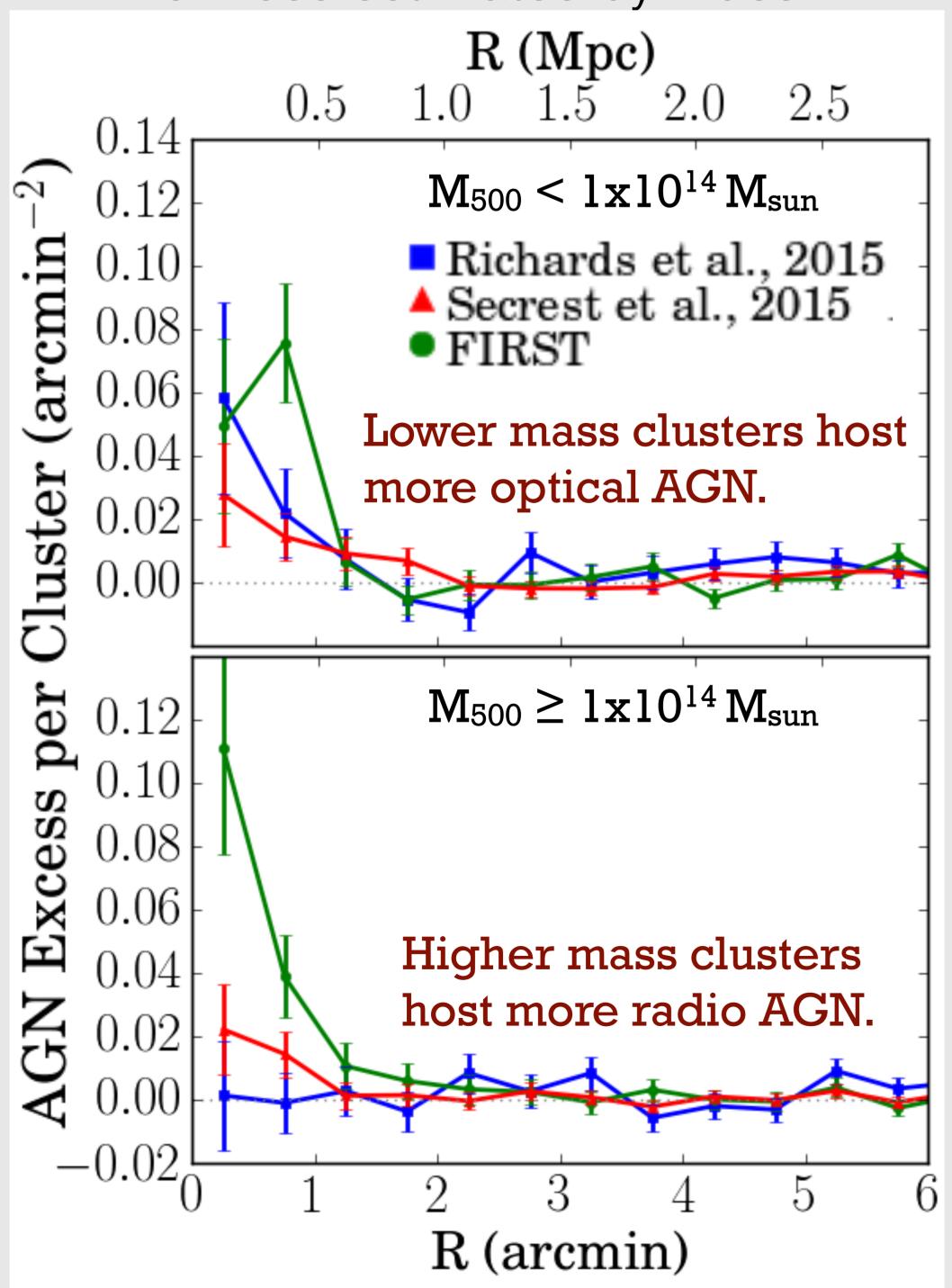
The cluster AGN distribution is determined by stacking upon cluster positions.

The background AGN counts is the average AGN surface density at 30'-60' from cluster center. Cluster centers determined via number-weighted algorithm using WISE data are accurate to within ~21".

## Results: Radial Distribution of AGN Excess in MaDCoWS Galaxy Clusters

### Cluster Mass

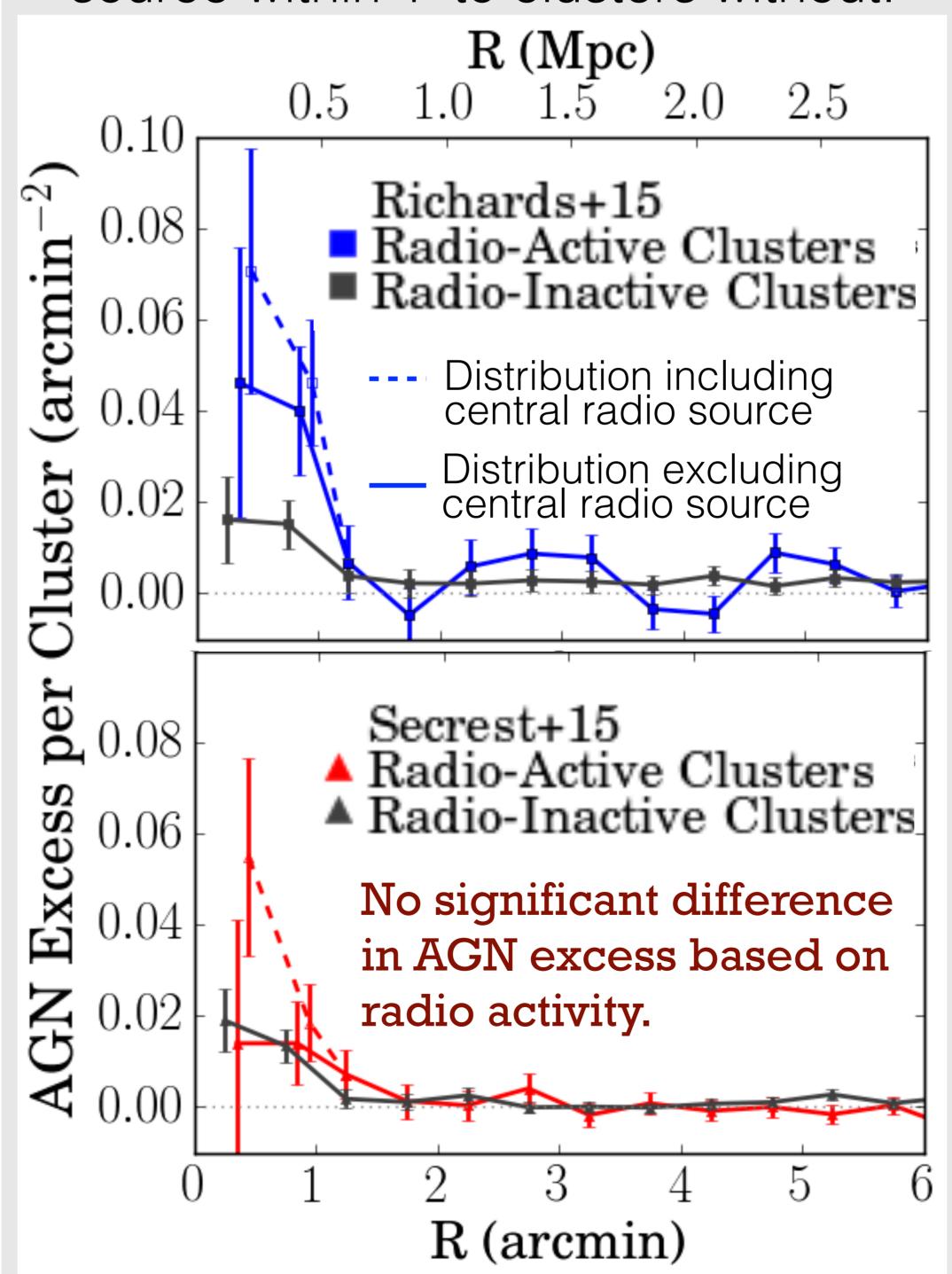
Divide 507 galaxy clusters with richness estimates by mass.



AGN Fueling via cold gas accretion more likely in low mass clusters. Radio-mode feedback important in high mass clusters.

## Central Radio Activity

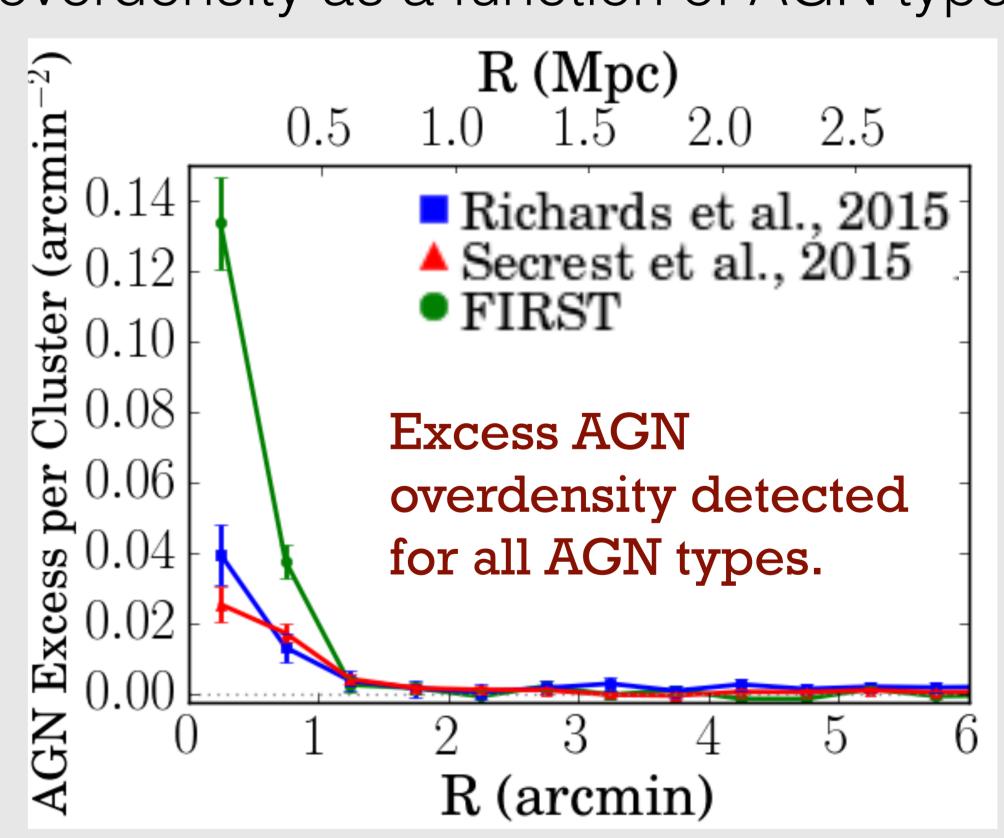
Compare 216 clusters with a radio source within 1' to clusters without.



Radio-mode feedback from central cluster radio source does not significantly impact the rate of AGN triggering.

## **AGN** Selection

Study the relative amplitudes of AGN overdensity as a function of AGN type.



Radio-selected AGN more concentrated towards cluster center than optical or mid-IR selected AGN.

#### References

Becker, R. H., White, R. L., & Helfand, D. J. 1995, ApJ, 450, 559 Richards, G. T., Myers, A. D., Peters, C. M., et al. 2015, ApJS,219, 39 Secrest, N. J., Dudik, R. P., Dorland, B. N., et al. 2015, ApJS,221, 12





MaDCoWs